Spring 2015 CCNY, Astronomy 30500, class # 16174, section ST

Tuesday 5-6:20 Pl NAC 1/2 ⁻	M 5-7	ursdays 7 PM R 3	Topics	
Jan Feb. 3 10		29 5 	Celestial observations & explanations 1 Celestial observations & explanations 2	
17 24		19 26	Celestial observations & explanations 3 Celestial observations & explanations 4	
March 3	test	\rightarrow Test #1 (30% of grade)		
10 17 24 April 31		5 12 19 26 2	Early history, Kepler's Laws, Galileo Newton's Laws, Gravitation Planet survey Light, wave properties, E/M waves, spectra Inverse square law. Stars: m&Mand L	
14	test	→ Test #2 (30% of grade)		
21 May 28 5 12	7	16 23 30 14	The sun H-R diagram; star classes, life, death Galaxies, Distance measurement Hubble's Law, Big Bang, Cosmology Life in the universe? Exoplanets.	
		21 (6 – 8:15 PM)	\rightarrow Final exam (40% of grade)	

• Astronomy is fascinating, beautiful, and in the news. Regardless of their beliefs and philosophy, intelligent people must consider where we are in the vast universe and what's happening "out there." Hope you'll learn a lot, enjoy the view, and grow from it. This class will definitely change and expand your world-view. You'll see things differently. Welcome.

• Text: Explorations. An Introduction to Astronomy, 7E, Arny & Scheider, 2012, McGraw-Hill

• A calculator is needed. Some math does appear. Inexpensive is fine. If it has an X^Y or Y^X button it has everything else you'll need for most science, math, and business classes.

• Lecture is meant to introduce the textbook readings. Even the book's photographs, diagrams, graphs, and math examples are to be studied. <u>Many test questions are based on book readings</u> <u>besides lecture notes.</u> The lecture is generally via a Power Point presentation which is posted onto BlackBoard just before the lecture so you can preview it and review it.

• **BlackBoard is vital!** It's used for announcements, Power Point lectures, downloads, and your grade record. You must learn how to enter it and to navigate within it.

• There will be three tests and an <u>optional</u> extra credit report (only 1). Extra credit projects will be described in a BlackBoard file. There's no guarantee but if done well (up to +4 points) it <u>might</u> move your final grade up by a half-letter, like C- \rightarrow C, C+ \rightarrow B-, B \rightarrow B+. But just going to an event or

planetarium show is not enough. A report is required, as described in the instructions. Write the report promptly after a visit or show for full credit.

• DON'T BE ABSENT. DON'T BE LATE. Attendance will be taken. Absences will not subtract from your average but will determine how much of a curve (0 to 9 points) will be added to your test average. More than 9 absences will result in no curve being given to you so *you'll sink or swim on just your raw test average*.

• Absent from a test? Notify me immediately by email as to why. Be ready for a make-up test during the next class. Bring a note. Otherwise, a 0 will be given.

• A photo ID will be required during tests.

• Review sheets for tests will be put in BlackBoard. You may use your answers during the tests. HOWEVER, all your answers and other notes must be compressed onto one sheet of paper and must <u>handwritten</u>. You cannot use photocopies from the textbook, nor of PowerPoint lecture slides, nor of someone else's work. Only notes <u>you wrote by hand</u> from the review sheet and readings can be used, and just one sheet (both sides OK.)

• Readings from the text will be assigned on separate BlackBoard announcements or on the last slide of each PowerPoint presentation. Read the pages assigned and try some questions at the chapter's end to cement your understanding. Homework will not be collected. The wise student will do it!

• The Final is not comprehensive. It only covers topics taught or assigned after Test 2.

General Education Information

As part of the College's General Education Curriculum, this course is designed to enhance your understanding of artistic issues and how they are studied. Students successfully completing this course will develop the following proficiencies:

Oral and written communication skills - Students will produce well-reasoned written or oral arguments using evidence to support conclusions.

Critical thinking skills - Students will evaluate evidence and arguments critically or analytically.

Information literacy - Students will gather, interpret, and assess information from a variety of sources and points of view.

Scientific World proficiency – Students will identify and apply the fundamental concepts and methods of a discipline or interdisciplinary field exploring the scientific world.

Detailed Astronomy Topics for Entire Course, Prof. Kiefer

Some topics and their Power Point presentations may assigned as independent study.

<u>Celestial Observations with Modern Explanations: 1</u> Course introduction Hierarchy of objects beyond Earth Powers of Ten- relative sizes of all things Rotation- day/night, 15°/h Polaris and latitude Constellations- 88 official, asterisms Revolution- changing constellations defines the year Precession- cycle of North stars and equinox point Planets- 5 ancient "wanderers", an orrery Major and minor planets, "demotion" of Pluto Retrograde motion- major problem for ancients The ecliptic, Zodiac constellations, nebular hypothesis

<u>Celestial Observations with Modern Explanations: 2</u> Moon, moon phases Eclipses Tides- daily and weekly variations explained Changing size of sun's disk, elliptical orbits Why one side of moon always faces Earth, tidal locking, 1 rotation = 1 revolution, other examples Earthshine

Celestial Observations with Modern Explanations: 3

Key measurements for astronomical distances: miles, km, m, AU, ly, pc Reason for the seasons "Equation" for summer Inventory of the Solar System Asteroids Meteoroids, meteors, meteorites Comets, elliptical/hyperbolic orbits Relative sizes of Solar System objects

<u>Celestial Observations with Modern Explanations: 4</u> The milky way and The Milky Way Galaxy Proper motion of stars

Nebulae: •gaseous (emission/reflection) •star clusters (open/globular) •planetary nebulae around exploded stars •other galaxies

Brief History of Western Astronomy Mystical, magical, and mechanical views Geocentric model (Ptolemaic) Heliocentric model (Copernican) Brahe Kepler, three laws of planetary motion, math problems involving Galileo's five central discoveries: moon features, sunspots, milky way, moons of Jupiter, phases of Venus and Mercury

<u>Newton's Laws (physics in astronomy)</u> Three laws of force and motion Mathematics of 2^{nd} Law, F = maUniversal gravitation, F_G equation Inverse square relationship G and g, mass and weight Circular motion, centripetal force F_C Understanding revolving in orbits

Planets

Key facts and unique features Temperature; absolute zero; °C, °F, K scales

Telescopes

Refractors/reflectors Curved mirrors and lenses, focal length (f) Three functions: gather light, resolve, magnify Importance of front diameter and eyepiece f length Key aspects in buying an amateur telescope

<u>Light</u>

Four representations: waves, ripples, rays, particles Standard aspects of all waves: medium, A, λ , f, T, v Problems using wave equation and relationship between λ , f, and v Light as an E/M wave. Color. Speed Dispersion of white light into colors with exact λ s by prisms and diffraction gratings E/M waves: gamma X, UV, visible, IR, micro, radio Inverse square law for light and stars Three measurements for star brightness: m, M, L Magnitude scale of star brightness Linking m, M, and d (distance) Linking M and L

Spectra

Spectroscopy, color "fingerprint" for each element Continuous, emission ("bright line"), and absorption ("dark line") Doppler effect in sound and in spectral lines Star spectra for composition, % composition, velocity, surface temperature, spin, pressure, magnetic field Detection of exoplanets by star oscillations

The sun

Stats about, special features, place among all stars Source of energy, nuclear fusion, review of atoms Hydrostatic equilibrium Major zones within and around it, solar wind

<u>Star Classification and the H-R diagram</u> Classifications: $A \rightarrow P$, now by O-B-A-F-G-K-M Classification, surface temperature, and visible color Parallax for near-by star distances, $pc = 1/p^{"}$ The H-R diagram, various axes possible Four star types: MS, RG, BG, WD Information compressed within an H-R graph

<u>Star birth and death</u> Formation of stars, mass is critical Why stars die. Subsequent fusions after H → He Upsets to hydrostatic equilibrium Mechanisms for star death: peter out to BD; explode to PN & WD, or SN & NS, or SN & BH Comparative lifetimes

Ladder rungs for distance measurement Radar/Transits of Venus Parallax Spectroscopic parallax- spectrum and H-R graph Cepheid-type variable stars, period-L link Tully-Fisher for galaxies, rotation-L link Type 1A SN with constant L (standard candle) Useful distance ranges for these methods

Galaxies and Hubble's Law1930s: vast, new view of universe's sizeRooted in Cepheid and 1A SN starsEdwin Hubble's impactGalaxy typesHubble's graph and equation, $v = H_0 x d$ Significance of H_0 slope of 70 km/s/MpcHubble's Law as final rung for distancesBig Bang

The Big BangNeither big nor bangCosmic inflation followed by cosmic expansionAnalogies for expansion of space itself carrying
galaxies apart $T_0 = 1/H_0 = 13.7$ Gy (billion years) agoContrast to Steady State viewEvidences:•Hubble expansion•Einstein's general theory of relativity•Uniform background radio waves,
Penzias and Wilson, WMAP satellite

- •Nucleosynthesis of elements 1-3, 4-26, 27-92
- •Furthest galaxy distances to date

Major mysteries being explored today Dark matter Dark energy Black holes Exoplanets Cosmology- past/future of universe, curvature of space, 4 fundamental forces Structure of matter and antimatter, Standard Model, newest research on Higgs particle Life in the universe- conditions, likely places, radio Communication Future of space exploration The moment of creation- sequences in Big Bang The mystery of existence: Why should anything exist? How can the universe emerge from nothing? Life in the Universe Possibilities and probabilities The Drake equation Exoplanets- number to date, "hot Jupiters" Radial velocity (spectrum) and transit methods Viable stars Habitable zone Conditions for evolution Urey-Miller experiment

- Extremophiles
- Broadcast technology
- How long do advanced civilizations last?

Spring 2015